# Visualization Tools for Adaptive Mesh Refinement Data

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# Outline

- Introduction to Berger-Colella AMR
- Visualization of Scalar AMR Data
- Specialized AMR Visualization Tools
- Visualization Tools with AMR Support
- Short overview of Vislt





## **Adaptive Mesh Refinement**

- Computational fluid dynamics technique
- Topological simplicity of regular grids
- Adaptivity of unstructured meshes
- Nested rectilinear patches, increasing resolution
  - Reduce simulation time
  - Reduce storage space
- Berger-Colella AMR: axis-aligned patches
- Very often: Cell centered data





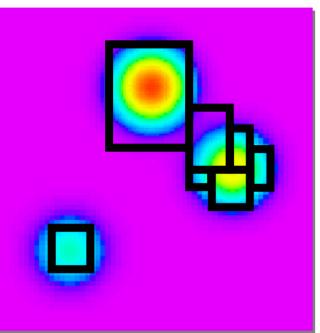
#### **Berger-Colella AMR Format**

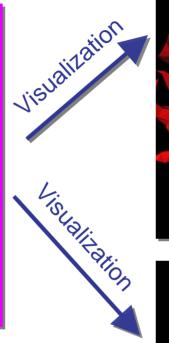
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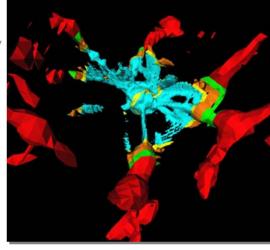




# **Effective Visualization of Scalar AMR Data**

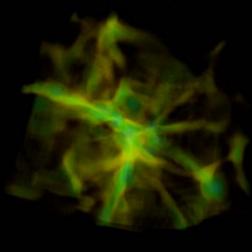






Isosurfaces

•Extraction of continuous "crack-free" isosurfaces



"Direct Volume Rendering"

Effective utilization of the hierarchy for efficient rendering
Good interpolation functions



Hierarchical AMR simulation

Aim: Use inherently hierarchical structure for efficient visualization





# AMR Visualization In the Beginning

- Translation of AMR to unstructured meshes [Norman et al. 1999]
  - Visualization with standard tool (VTK, IDL, AVS)
  - Ineffective utilization of computational resources
- Direct Volume Rendering
  - Mention AMR data without further details [Max 1993]
  - PARAMESH [Ma 1999]
    - Resampling
    - Block-based





# Isosurfaces

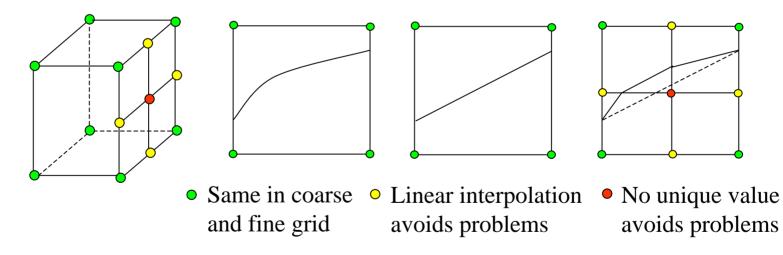






# **Marching Cubes and Dangling Nodes**

- Marching cubes needs vertex centered data
  - Resample data set to vertex centered case
- Dangling nodes := only present in fine level (yellow + red)
  - Choice of consistent values to avoid problems?



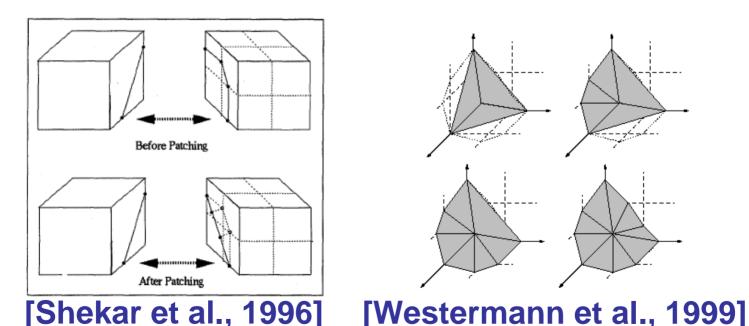
→ Compare [Westermann, Kobbelt, Ertl 1999]





## **Previous Crack-fixing Solutions**

- Mostly in context of Octree-based hierarchies
- [Shu et al., 1995]: Create polygon to fit crack
- [Shekhar et al., 1996]: Collapse polyline to line
- [Westermann et al., 1999]: Create triangle fan







# **First Approach: Use of Dual Grids**

- "Avoid interpolation whenever possible!"
- Avoid interpolation apart from linear interpolation along edges, which is part of marching cubes
- Use *dual grid* := grid formed by connecting cell centers







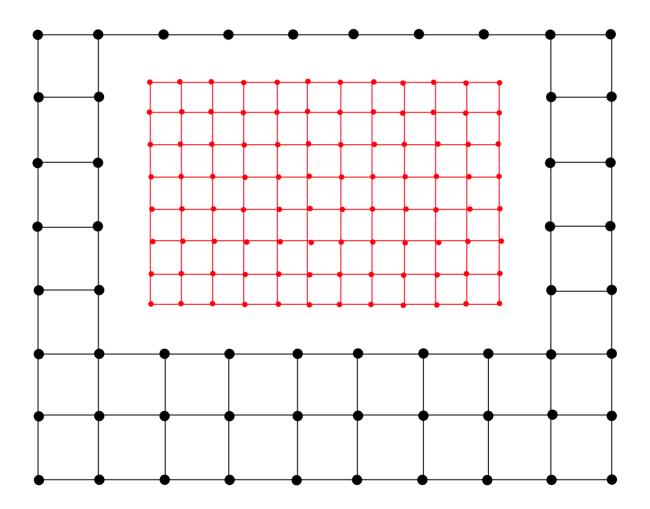
#### **Dual Grid – Original Grid**

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#### **Dual Grids**







# **Advantages of Dual Grid Approach**

- Use of values original data for marching cubes
- No dangling nodes
- Instead: Gaps between hierarchy levels!
- → Fill those gaps with *stitch cells*







# **Stitching the Gaps**

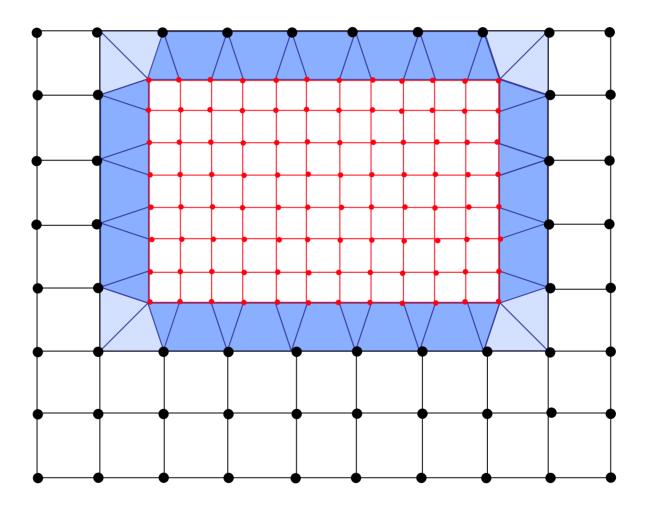
- Tessellation scheme for filling the gap between two hierarchy levels
- Constraints
  - Only gap region is tessellated
  - The complete gap region is tessellated
  - Only vertices, edges and *complete* faces are shared

➔ In 3D space: Cannot use tetrahedra because cells must share quadrilaterals as faces





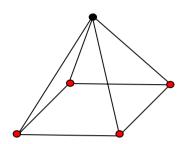
#### **Stitching Process**

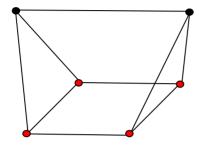


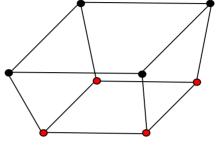




#### Stitch Cells – 3D Case

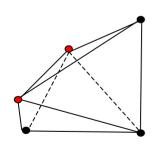


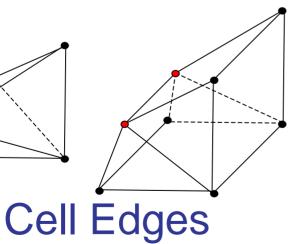


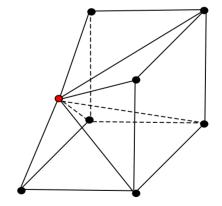




#### **Cell Faces**





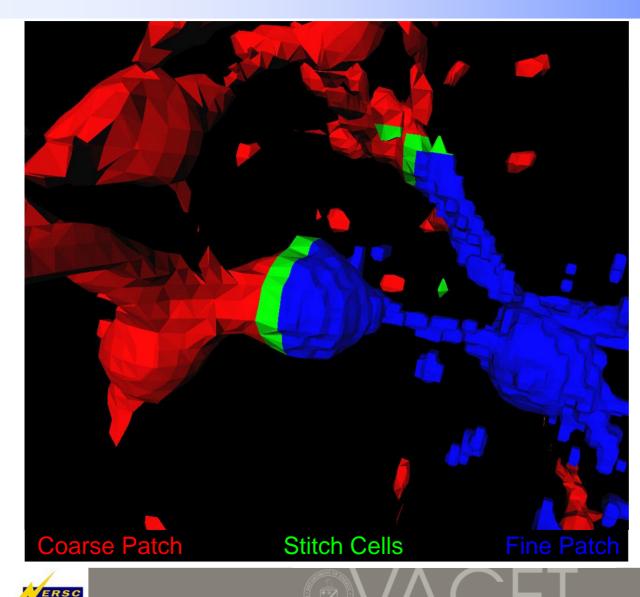


#### **Cell Vertices**





#### **First Results**



AMR simulation of star cluster formation

Root level 32x32x32

[Data set: Greg Bryan, Theoretical Astronomy Group, MIT]

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## **Multiple Patches**

- Multiple patches can be connected using the same scheme
- However: Special care must be taken with adjacent fine patches.
- Must "merge" adjacent grids (i. e., "upgrade" edges to quadrilaterals and vertices to edges)







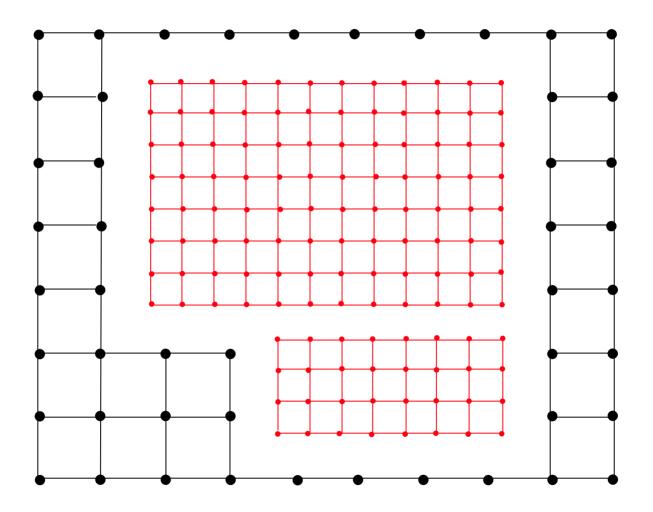
#### **Multiple Patches – Example**

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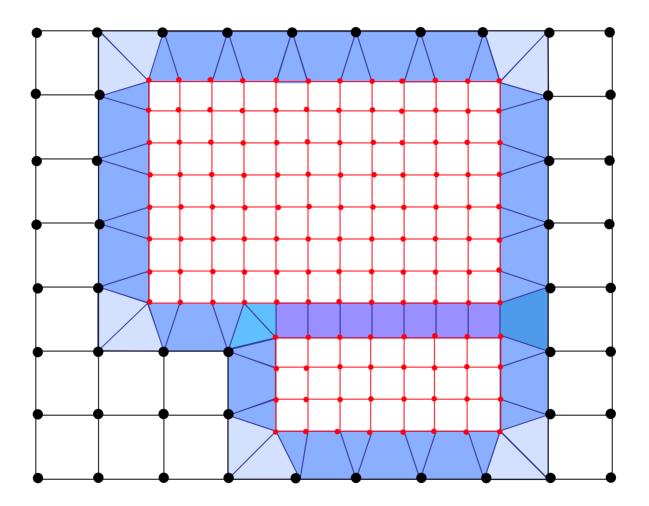
#### **Multiple Patches – Example**





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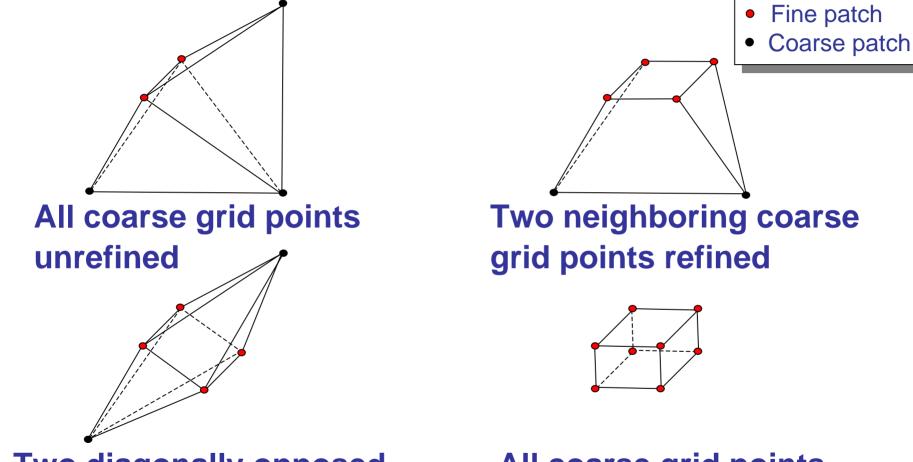
#### **Multiple Patches – Example**







#### Multiple Patches – Fine Edge to Coarse Edges



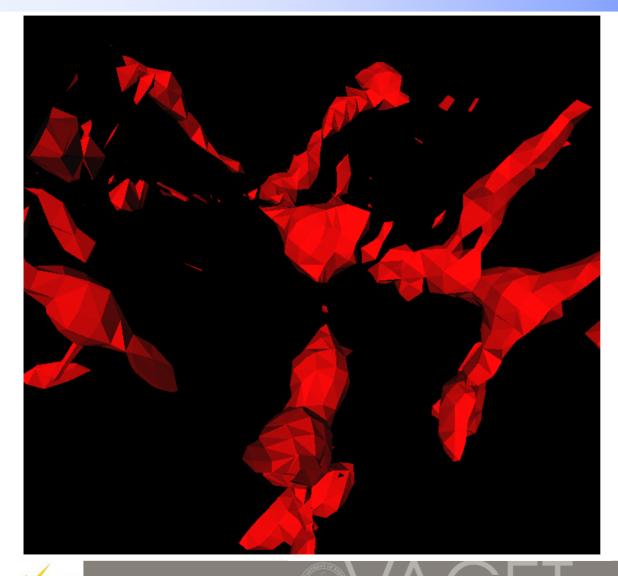
Two diagonally opposed coarse grid points refined

# All coarse grid points refined





#### **Isosurface - One Level**



ERSC

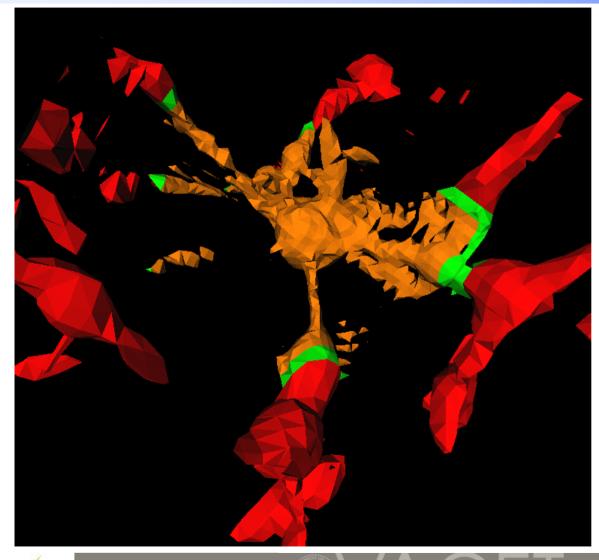
AMR simulation of star cluster formation

Root level 32x32x32

[Data set: Greg Bryan, Theoretical Astronomy Group, MIT]



#### **Isosurface - Two Levels**



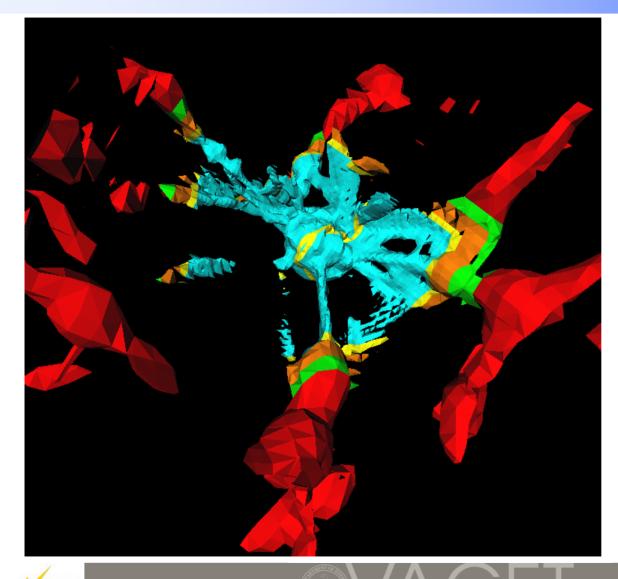
AMR simulation of star cluster formation

First level Stitch cells (1/2) Second level





#### **Isosurface - Three Levels**



ERSC

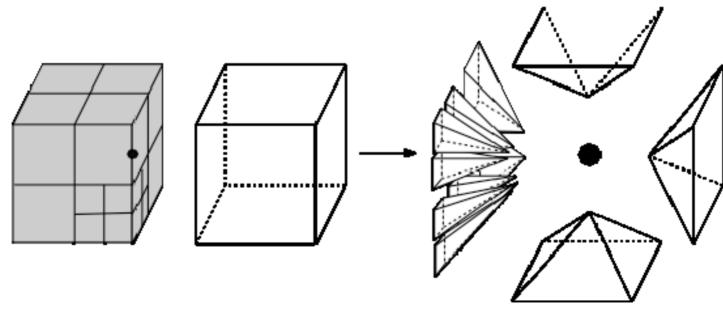
AMR simulation of star cluster formation

First level Stitch cells (1/2) Second level Stitch cells (2/3) Third level



#### Second Approach: Keep Grid

- Vertex/node centered data
- Retain "identity" of cells (debugging)
- Subdivide boundary cells into pyramids
  - Eliminates "non-linear" hanging nodes
  - Standard isosurface techniques for pyramids

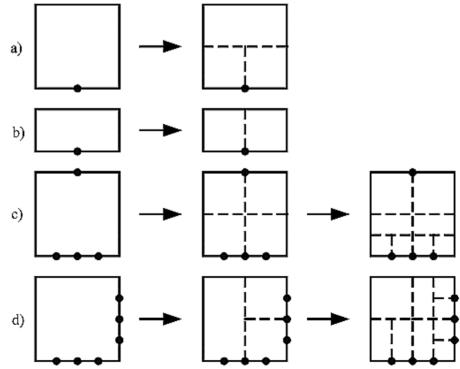






#### **2D Case**

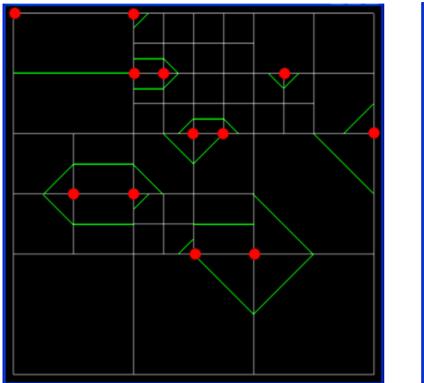
- Forms basis of 3D case
- Split cell faces to eliminate hanging nodes along edges
- Obtain values at newly created hanging by linear interpolation

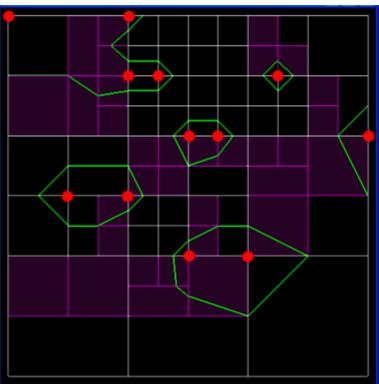






#### **2D Results**





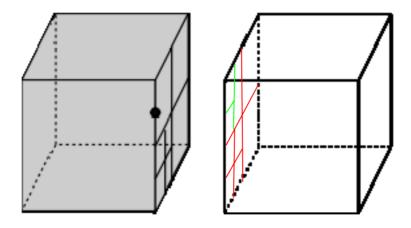






#### **3D Cell Face Subdivision**

- Subdivide lower-resolution cell face to match higher resolution face
- Subdivide cell face to eliminate hanging nodes

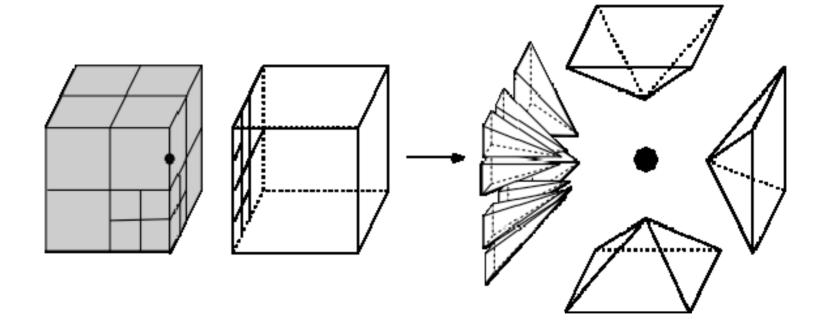






#### **3D Cell Subdivision**

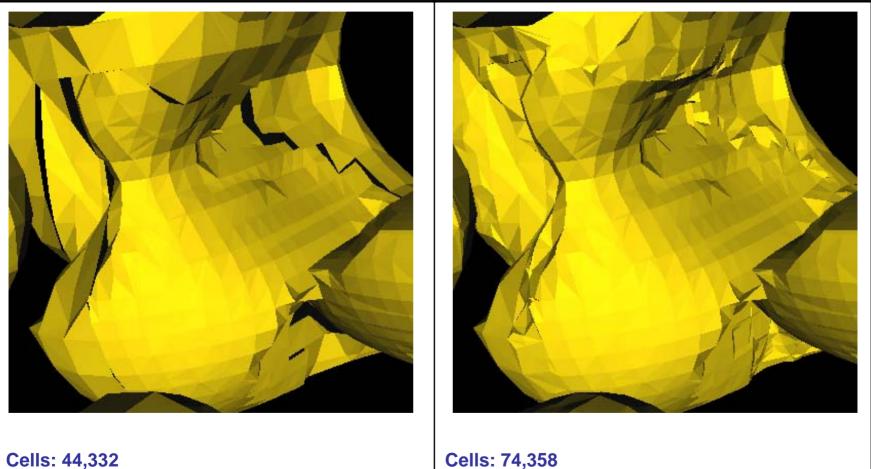
Subdivide cell into pyramids with common apex point







#### **Second Approach – Results**



Triangles: 10,456

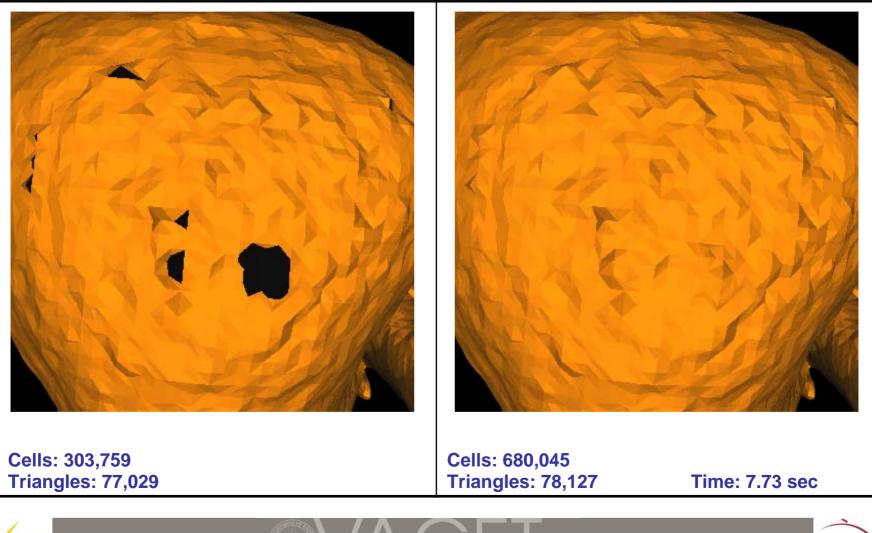
Cells: 74,358 Triangles: 14,332

Time: 2.30 sec





#### **Second Approach – Results**







# **Volume Rendering**





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#### Hardware-accelerated Preview of AMR Data

- Interactive DVR for choosing view point and transfer function
- Subdivide data set in regions of constant resolution
- → AMR Partition Tree (generalized kD-tree)
- Traverse "AMR Partition tree" and render regions using hardware-accelerated DVR





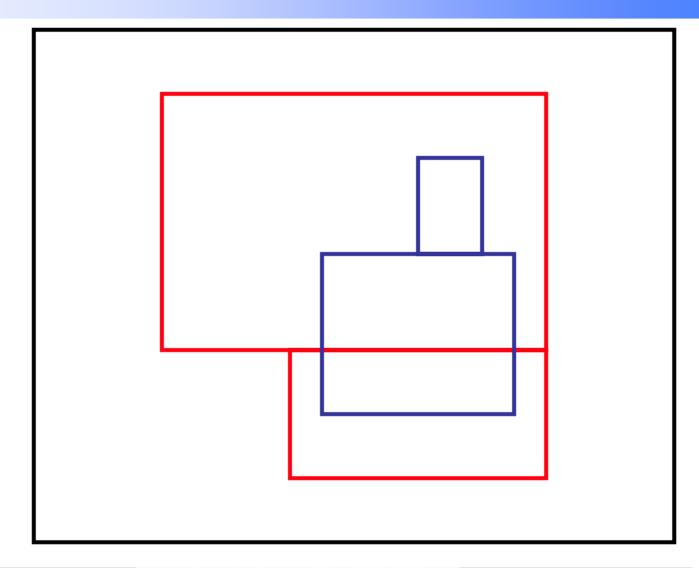
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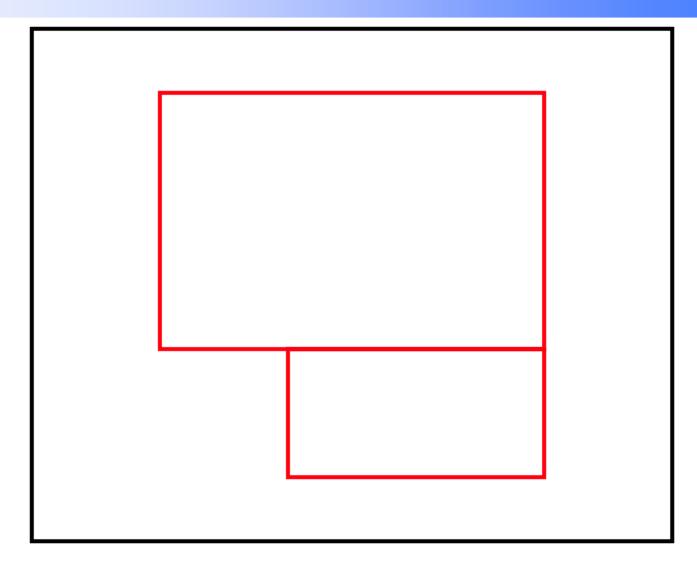
# Homogenization







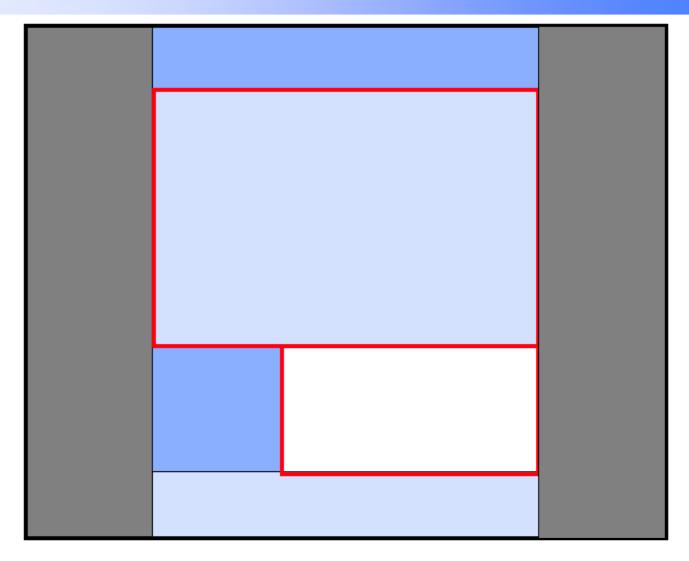
## Homogenization







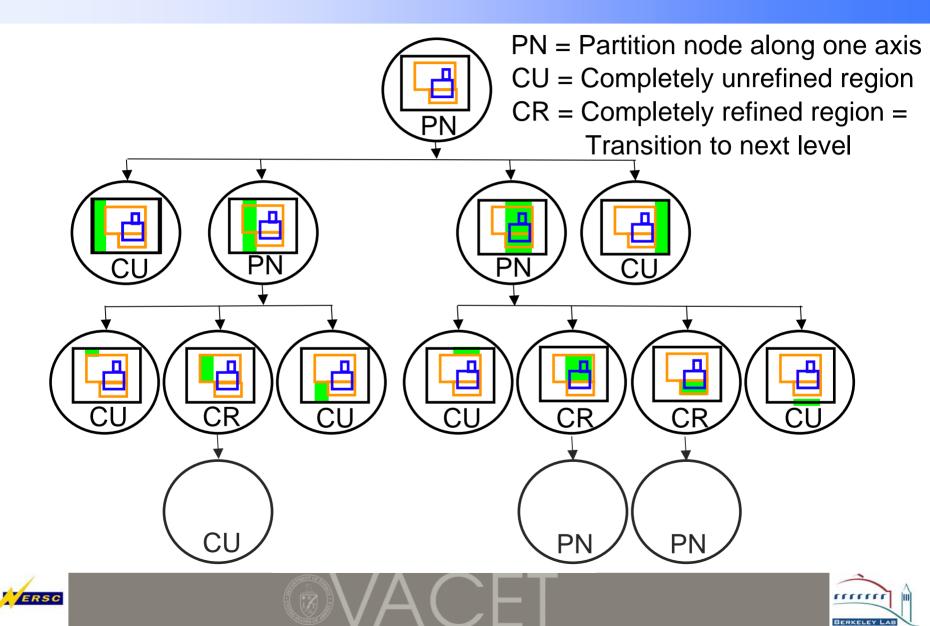
## Homogenization



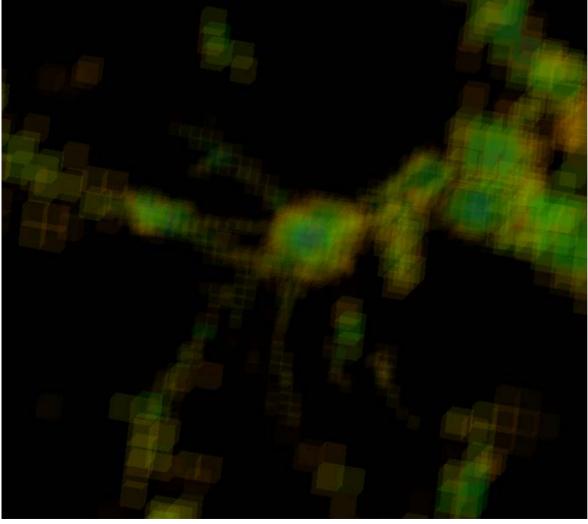




#### **Partition Tree – Example**



#### Hardware-accelerated Rendering – Interactive

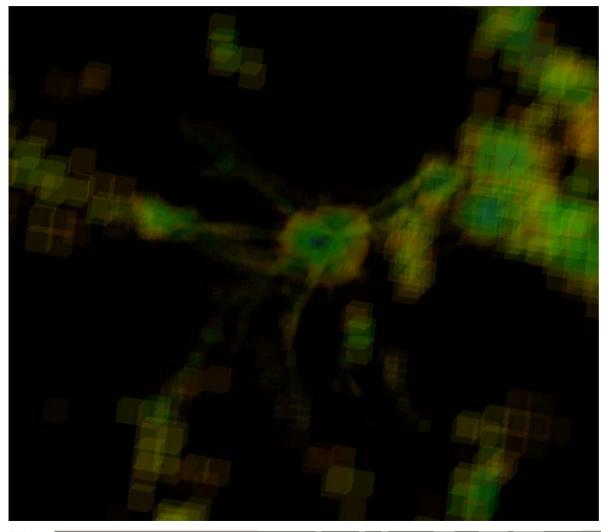


**AMR** simulation of star cluster formation **Root level** 32x32x32 [Data set: Greg Bryan, **Theoretical** Astronomy Group, MIT]





#### Hardware-accelerated Rendering – Maximum Quality



**AMR** simulation of star cluster formation **Root level** 32x32x32 [Data set: Greg Bryan, **Theoretical** 

Astronomy

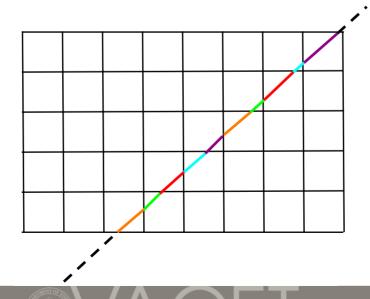
Group, MIT]





# **High-quality DVR of AMR Data**

- Use "cell projection" [Ma & Crockett 1997] to display individual patches
  - Traverse patches and construct ray segments [object space based]
  - Ma & Crockett: Sort ray segments

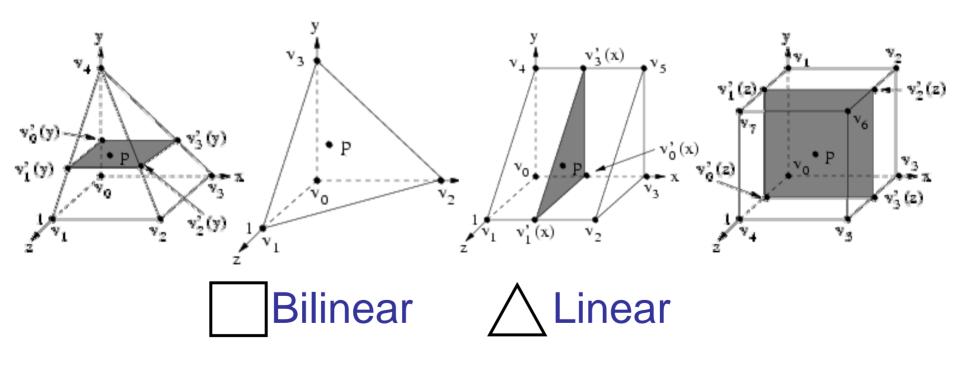






# Interpolation

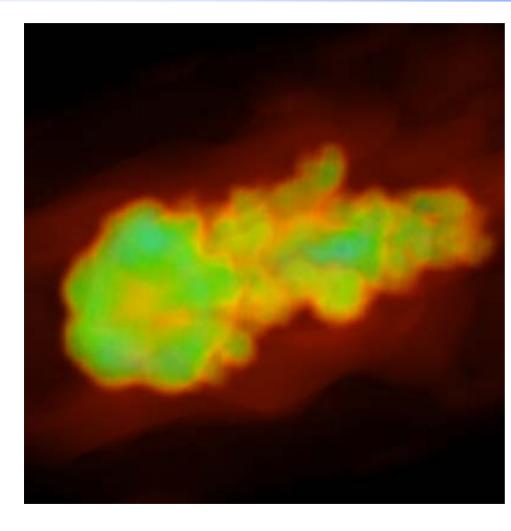
- Nearest neighbor (constant) interpolation → debugging
- Piecewise Linear Method (PLM) → Discontinuities
- Dual grids (trilinear) and stitch cells







#### Interpolation with Stitch Cells – One Hierarchy Level



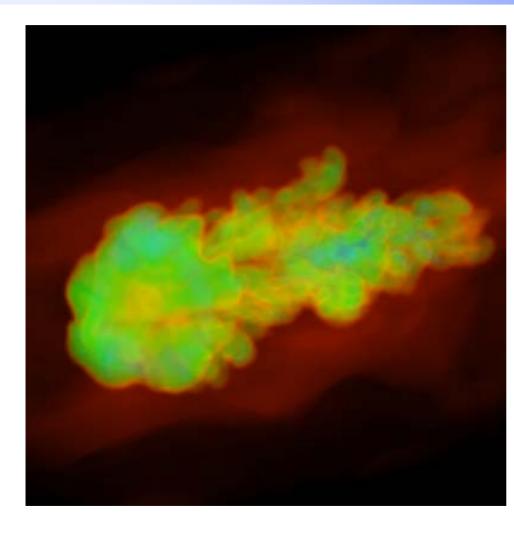
Simulation of an Argon bubble in a surrounding gas hit by a shockwave

[Data set: Center for Computational Sciences and Engineering (CCSE), Lawrence Berkeley National Laboratory]





#### Interpolation with Stitch Cells – Two Hierarchy Levels



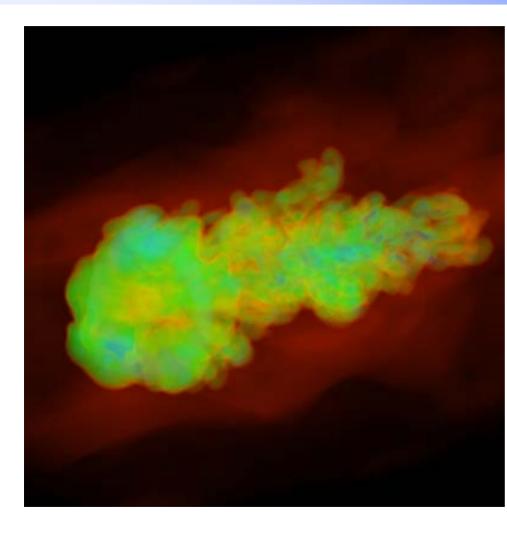
Simulation of an Argon bubble in a surrounding gas hit by a shockwave

[Data set: Center for Computational Sciences and Engineering (CCSE), Lawrence Berkeley National Laboratory]





#### Interpolation with Stitch Cells – Three Hierarchy Levels



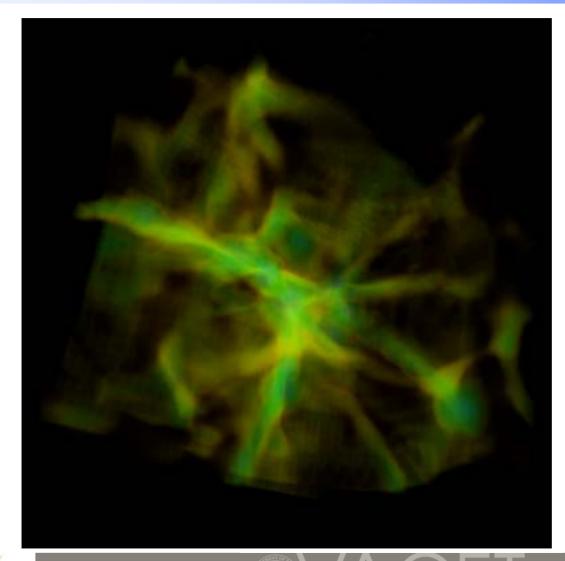
Simulation of an Argon bubble in a surrounding gas hit by a shockwave

[Data set: Center for Computational Sciences and Engineering (CCSE), Lawrence Berkeley National Laboratory]





### **No Transfer Function Scaling**



AMR simulation of star cluster formation

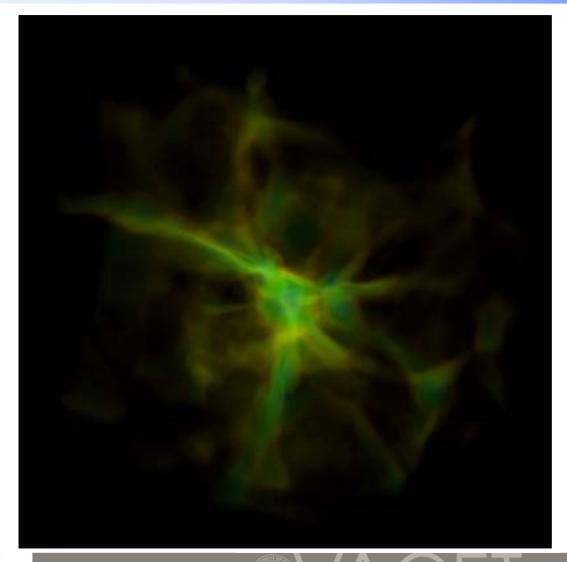
Root level 32x32x32

[Data set: Greg Bryan, Theoretical Astronomy Group, MIT]





# **Opacity Scaling**



AMR simulation of star cluster formation

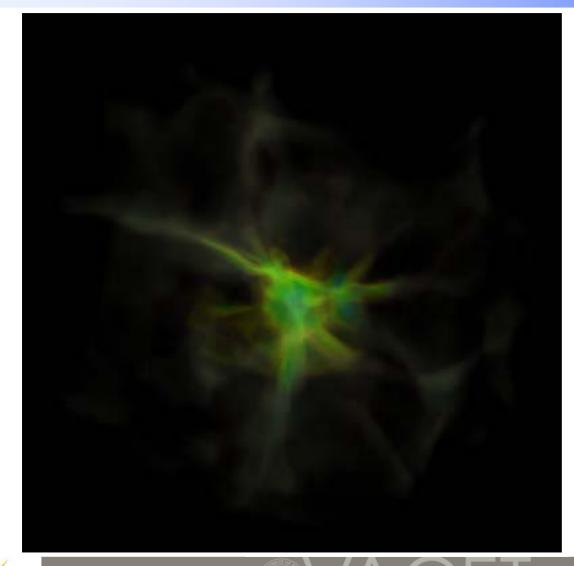
Root level 32x32x32

[Data set: Greg Bryan, Theoretical Astronomy Group, MIT]





## **Opacity and Saturation Scaling**



AMR simulation of star cluster formation

Root level 32x32x32

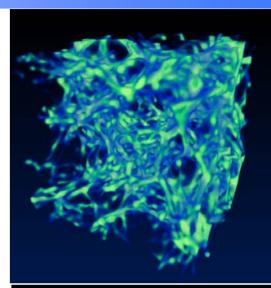
[Data set: Greg Bryan, Theoretical Astronomy Group, MIT]

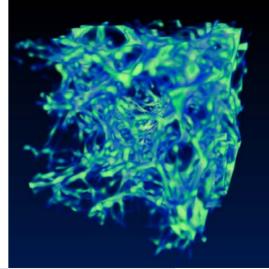




# **Texture-based AMR Volume Rendering**

- [Kähler & Hege, 2001 / 2002]
- Resample to node centered
- Subdivide in homogenous resolution regions (kD-tree)
- Minimize number of blocks using information about AMR grid placement algorithm
- Texture/Slicing-based volume rendering
- Optimized texture packing
- Adapt slice spacing & correct opacity









## **Rendering the First Star of the Universe**

- [Kähler et al., 2002]: Application to astrophysical data set
   Texture-based volume renderer, Virtual Director, CAVE
- Aired on Discovery Channel

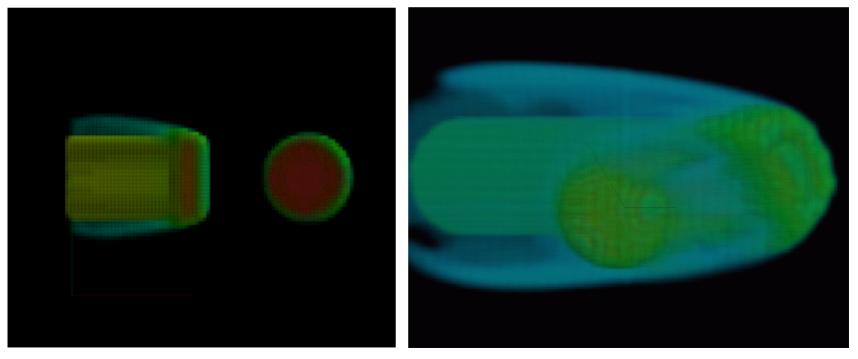






## Splatting-based Volume Rendering of AMR Data

- [Park et al., 2002]
- kD-tree- and Octree-based domain subdivision
- Specify isovalue range and transfer function
- Rendering using hierarchical splatting







### **Direct Volume Rendering of AMR Data**

- [Kreylos et al., 2002]
- Homogenization using kD-tree
- Distributed rendering using texture-based slicing
- Cost-range decomposition







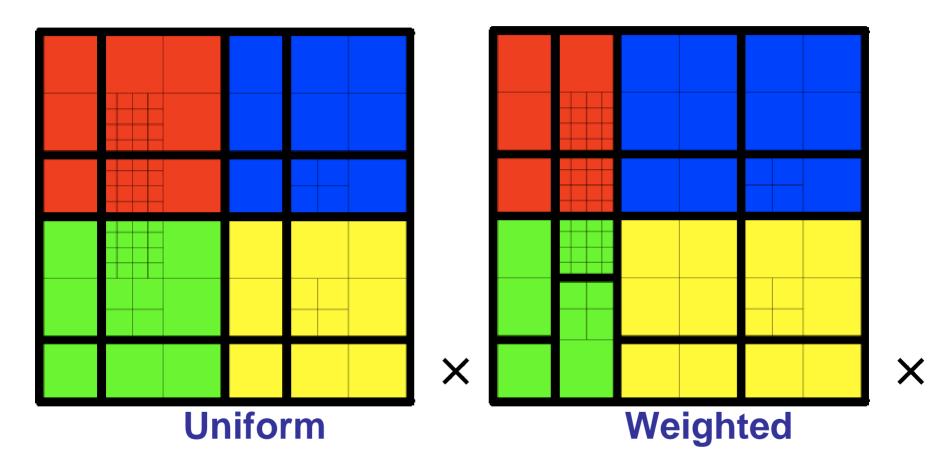
# **Framework for Parallel AMR Rendering**

- Efficient reimplementation of cell projection
  - Sort cells [Williams, Max & Stein 1998]
- Subdivision in object space with kD-tree
- Subdivision of first hierarchy level
  - Uniform: Blocks of approximately equal size
  - Weighted: Blocks of similar computational effort
- Subdivision in blocks of constant resolution
  - Unweighted
  - Weighted





#### Subdivision Strategies – Subdivision of the First Hierarchy Level

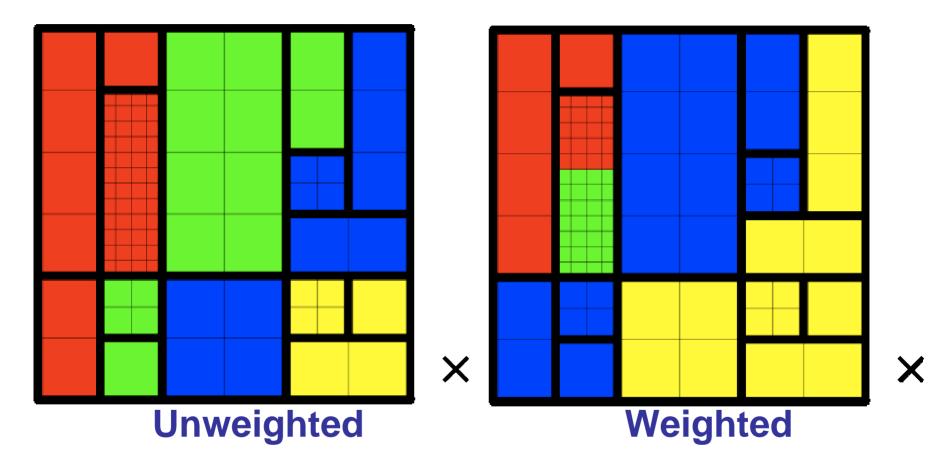


X = Viewpoint; Color = Assigned processor





## **Subdivision Strategies – Homogenization**

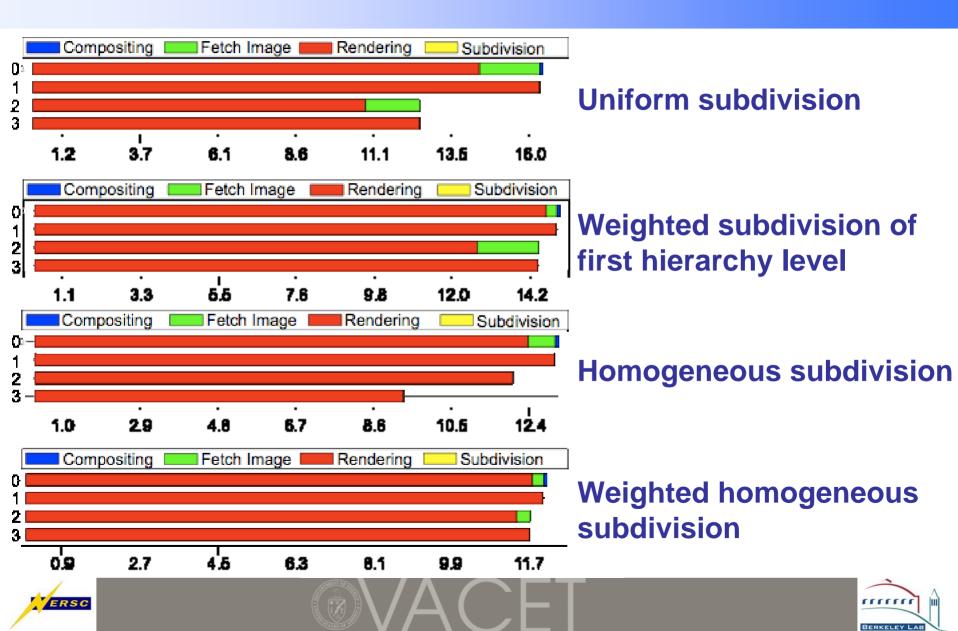


X = Viewpoint; Color = Assigned processor





# **Timing Results**



#### **Observations**

- Homogenization most efficient way to render AMR hierarchies
  - Computationally efficient
  - Use of standard methods
- Use of kD-tree currently standard way of describing subdivision
- Reasonable estimate of computational costs for rendering grid parts possible





## **GPU-Assisted Raycasting of AMR Data**

- [Kähler et al., 2006]
- Use raycasting instead of texture slicing
- Higher quality (improved precision, avoid varying sample distances)
- Sophisticated light model with wavelength dependent absorption







## **Visualization of Time-varying AMR Data**

- Feature-tracking
  - [Chen et al., 2003]
  - Isosurface visualization
  - Track connected components through time and AMR levels
- Remote visualization of time-dependent AMR data
  - [Kähler et al., 2005]
  - Interpolation scheme for "in-betweening" of hierarchy levels evolving at different simulation rates
  - Access remote simulation over network





# **Specialized Tools for AMR Data**

- ChomboVis
  - LBNL Applied Numerical Algorithms Group
  - Slicing and spreadsheets
  - Isosurfaces (w/ cracks)
  - Streamline computation (unpublished)
- AMR Vis
  - LBNL Center for Computational Sciences and Engineering
  - Shear-warp volume rendering (re-sampling)
  - Slicing and spreadsheets
  - Streamlines





# **Spreadsheets**

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- 98: multi_rect2d.silo	0.0 X Axis (cm)	(cm)	j=9	2.001755	2.107659			6 2.461094	2.585965		
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			<u>j=3</u>	1.688307	1.750399		1.875965		1.996760	2.052061	
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# **Visualization Tools with AMR Support**

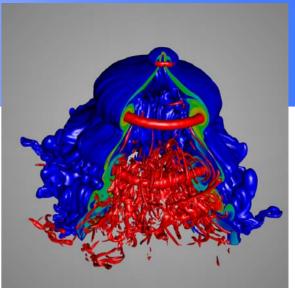
- ParaView
  - Support for reading AMR data sets (e.g., VTM)
  - Slicing, Isosurfaces (with cracks)
  - Volume rendering in development (commercial version)
- Amira
  - Some AMR support in internal collaboration version
  - Mainly volume rendering
- Vislt
  - Support for reading AMR data sets (e.g., Enzo, Boxlib, Chombo)
  - Wide range of visualizations including volume rendering, slices, isosurfaces (currently w/ cracks)



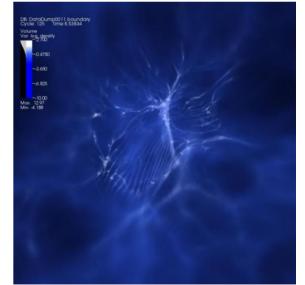




- Richly featured visualization and analysis tool for large data sets
- Data-parallel client server model, distribution on per patch-basis
- Use of meta-data / contracts to reduce amount of processed data
- Built for 5 use cases:
  - Data exploration
  - Visual debugging
  - Quantitative analysis
  - Presentation graphics
  - Comparative analysis



[Argon bubble subjected to shock Jeff Greenbough, LLNL]



[Logarithm of gas/dust density in Enzo star/galaxy simulation, Tom Abel & Matthew Turk, Kavli Institute]





# Vislt and AMR Data

- Supported as "first-class" data type
- Handled via "ghost-cells": Coarse cells that are refined are marked "ghost" in the lower level
- Isocontouring via resampling, cracks possible at level boundaries
- Work on rectilinear grids and skip ghost cells or "remove" results produced in ghost cells later on
- AMR capabilities currently under rapid development (planned as ChomboVis replacement this FY)
- http://www.llnl.gov/visit





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- AG Graphische Datenverarbeitung und Computergeometrie
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- National Science Foundation
- Office of Naval Research
- Army Research Office
- NASA Ames Research Center
- North Atlantic Treaty Organization
- ALSTOM Schilling Robotics, Chevron, General Atomics, Silicon Graphics, and ST Microelectronics, Inc.
- Stiftung für Innovation des Landes Rheinland-Pfalz





# **Questions?**





